



KIDS

KILO DEGREE SURVEY

**F. KÖHLINGER, B. JOACHIMI, S. JOUDAKI, L. MILLER
ON BEHALF OF THE COLLABORATION**

SCLSS WORKSHOP, OXFORD, 04-2018

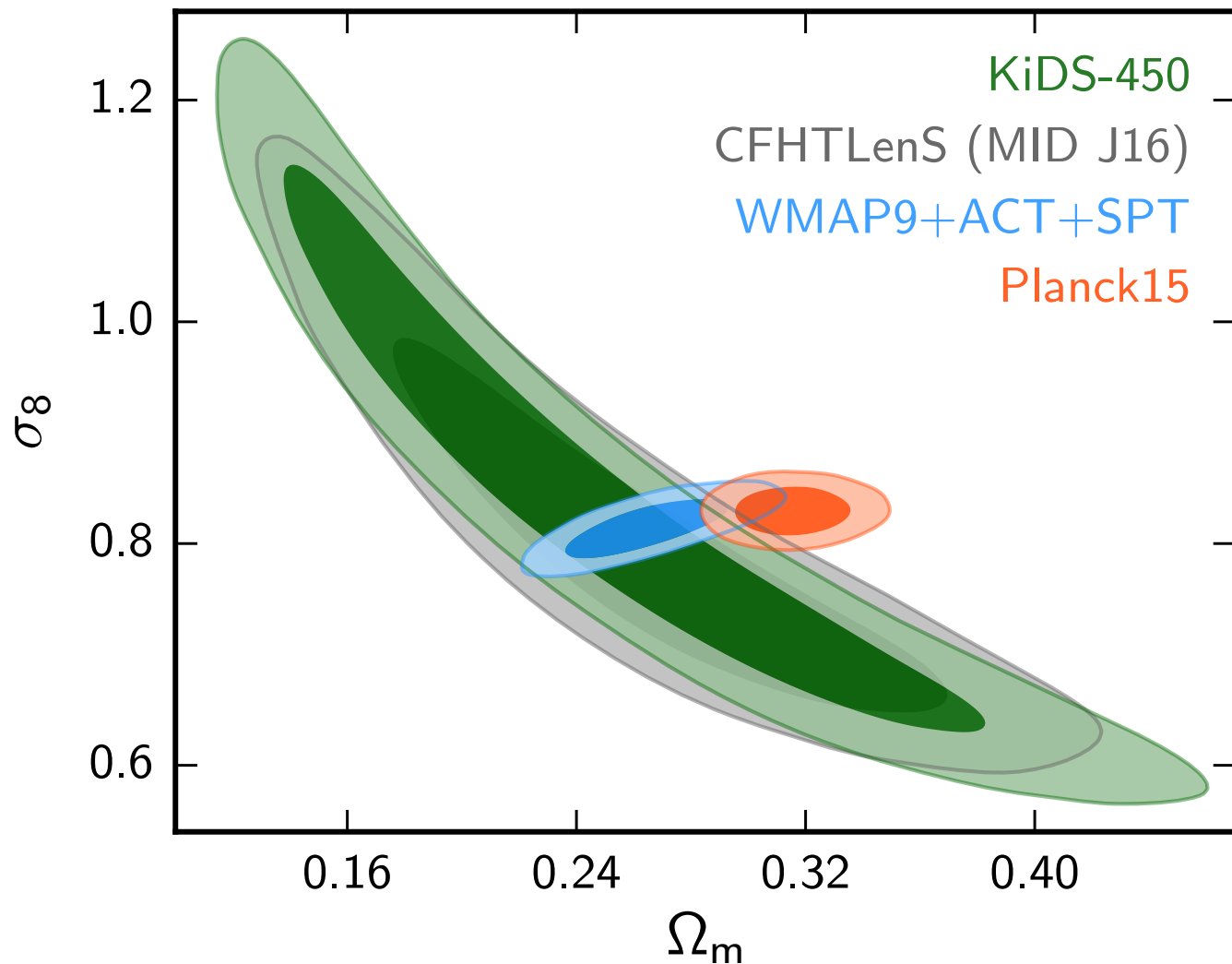
KiDS & DES survey properties

- Mirror/Focus (2.6 m Cassegrain vs 4.0 m Prime)
- Area (450 sq vs 1800 sq deg) —> 1350 vs 5000 sq deg
- Depth ($r \sim 24$ vs $r \sim 23$)
- Seeing (0.68 vs 0.96 arcsec)
- FOV (1.0 vs 3.0 deg²)
- Source density (~ 8.5 vs ~ 5.5 gal/arcmin²)
- Filters (ugri{ZYJHK_s} vs griz{Y})
- Team size (~ 30 vs ~ 130)

Differences in the analyses

- Matter power spectrum w/ baryonic feedback (HMCODE vs Halofit+cuts)
- Shear measurement (lensfit vs Metacalibration/im3shape)
- Photometric redshift calibration (spec-z vs 30-band photo-z)
- Photo-z uncertainty (bootstrap realizations vs mean-z shifts)
- Intrinsic galaxy alignments (A vs $\{A, \eta\}$)
- Covariance (analytic & numerical simulations)
- Propagation of shear calibration uncertainty (covariance vs free parameters)
- Lens sample (overlapping spec-z surveys vs RedMagic) \rightarrow RSDs

KiDS-450 cosmic shear

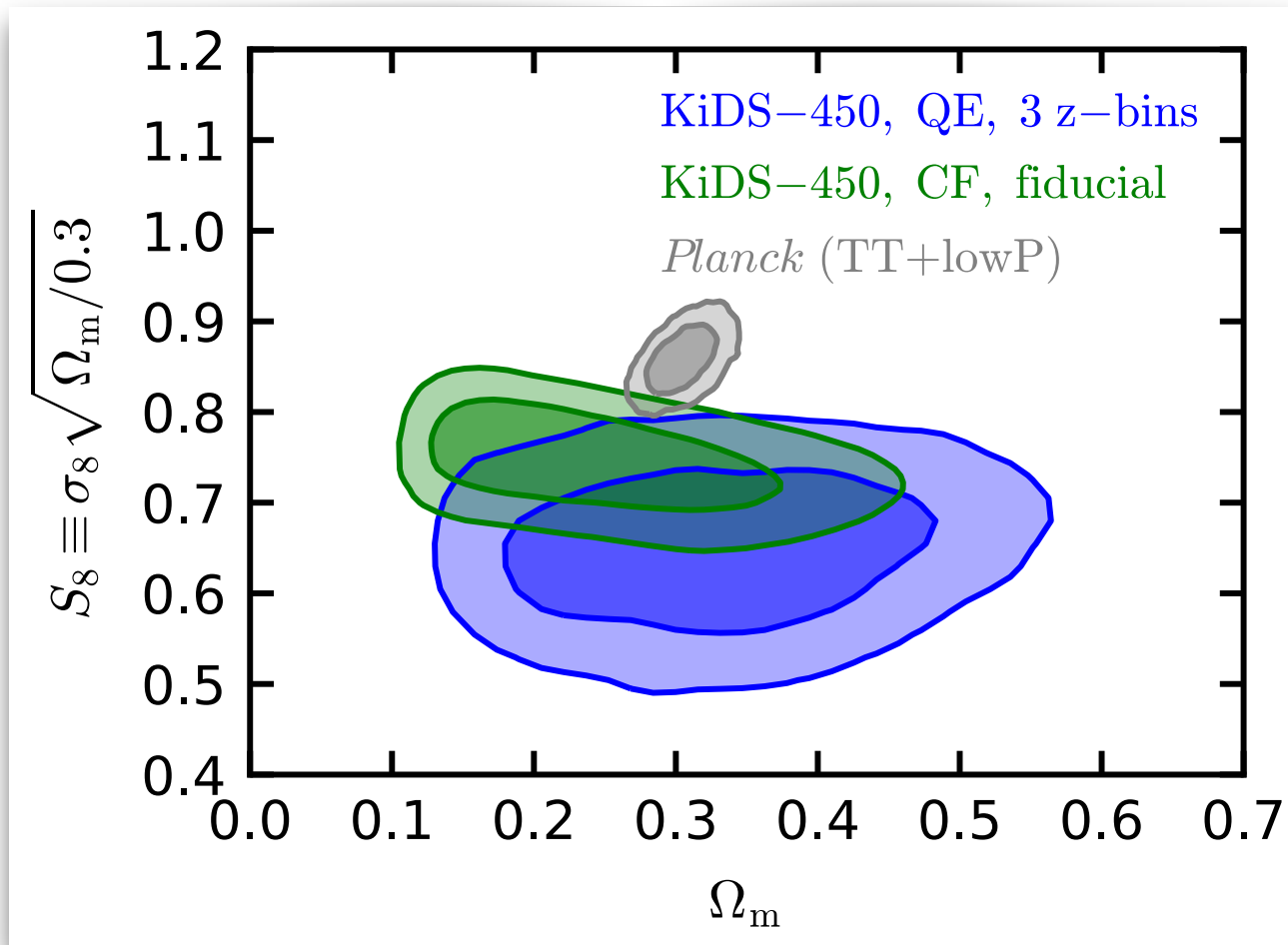


2.3 σ tension
with Planck -
unaccounted
systematics or
new physics?

Explored
extended
systematics and
cosmologies:
evolving DE

HILDEBRANDT ET AL
2017

KiDS-450 Optimal QE

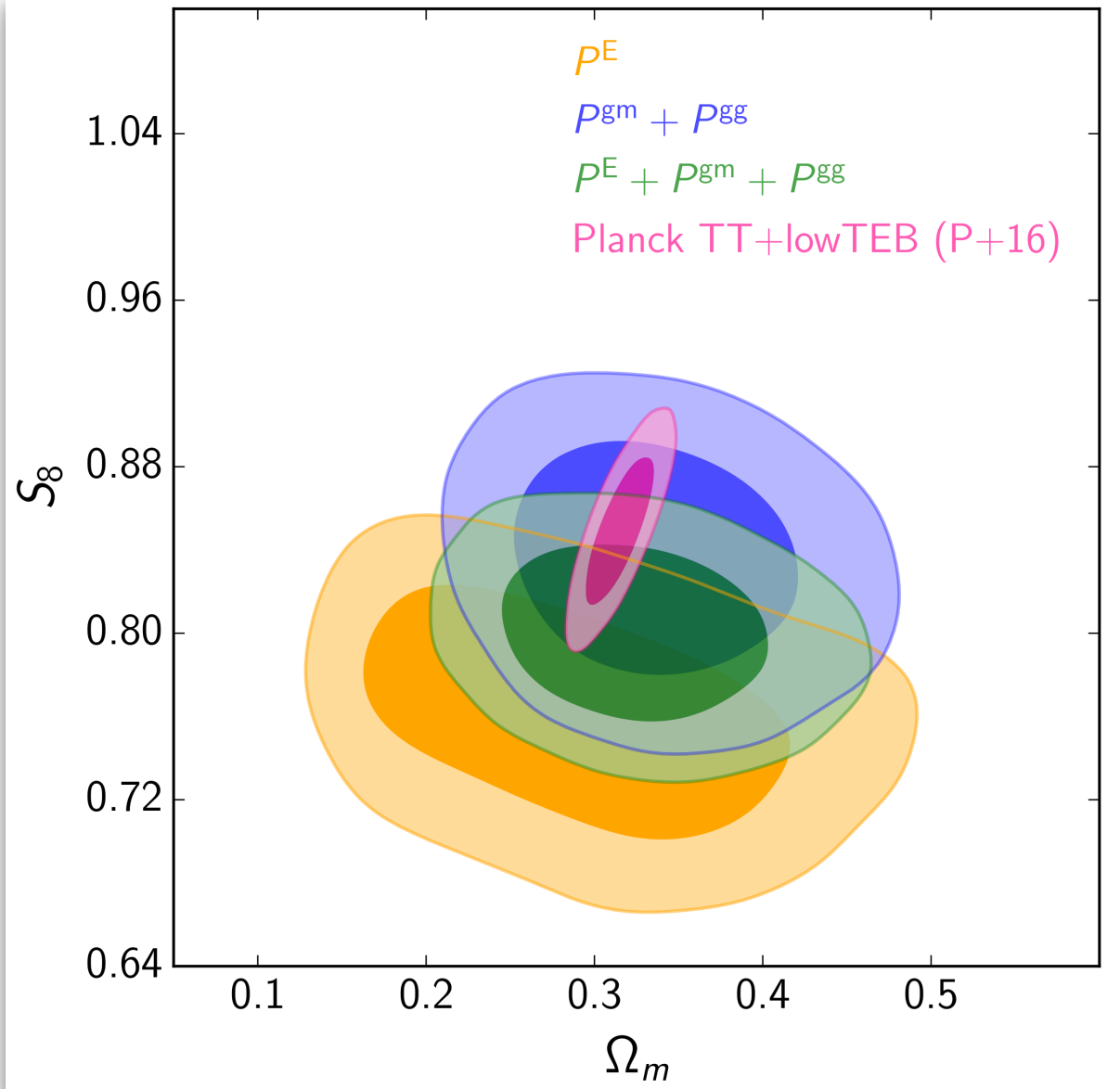


Fourier space analysis, less sensitive to small scales, favoring even lower S_8 .

In agreement with H17, particularly when CF analysis restricted to large scales.

KÖHLINGER ET AL
2017

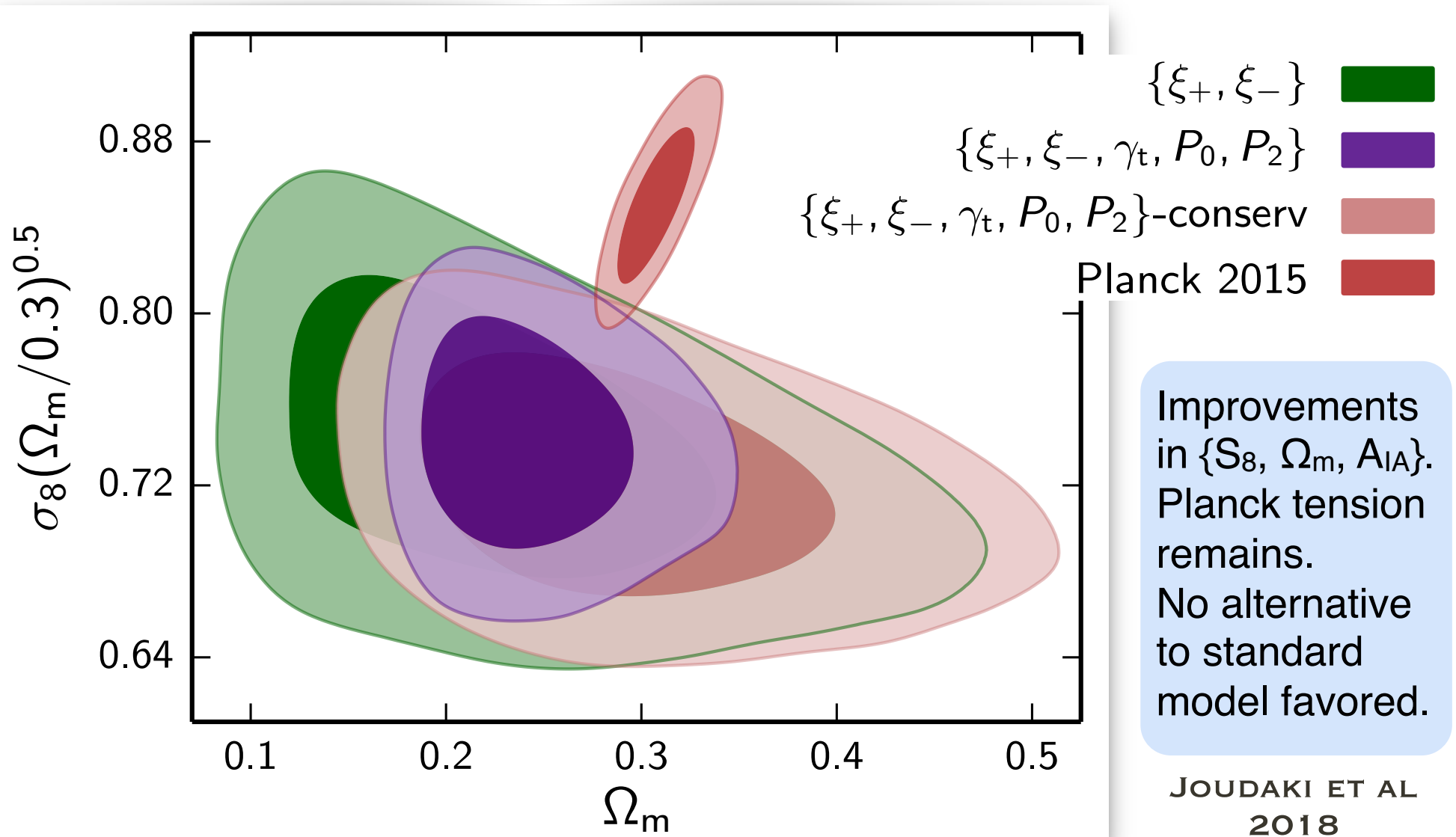
KiDS + GAMA (3x2pt)



Power spectrum analysis — cosmic shear P^E in agreement with H17. Larger S_8 preferred by GAMA clustering.

Combined probes improvements in $\{S_8, \Omega_m, A_{IA}\}$, in agreement with Planck and H17.

KiDS + 2dFLenS (3x2pt with RSDs)



K900/1350 opportunities & challenges

- Double/triple the area: improved statistics
- Photometric redshifts (will improve with VIKING overlap, 5th bin)
- Increase in spectroscopic overlap → systematics calibration & new physics
- Cross-survey measurement comparison: Lensing without borders

- Intrinsic alignments (nonlinear scales - A , L , z dependence, red/blue split)
- Matter power spectrum (current DM-only calibration ~5% level)
- Baryonic feedback (large spread in hydrodynamical simulations)
- Shear calibration (additive and multiplicative biases)
- Modeling of galaxy bias (linear/nonlinear, validate against simulations)
- Psychological systematics: blinding

Thanks for listening.